
Plenary Lecture | Plenary lecture

Plenary lecture V

Fri. Jul 19, 2024 9:15 AM - 10:15 AM Room A

[P5-01] Introduction, history, and applications of all speed Riemann flux schemes

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Keywords: scheme, all-speed Riemann flux, JAXA's research

Introduction, history, and applications of all-speed Riemann flux schemes

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1

Outline

1. Background and overview
2. Introduction of JAXA and the applications of SLAU type schemes in JAXA's research projects
3. History and theories of SLAU type scheme
 1. What is the Riemann Flux?
 2. Toward the AUSM type schemes
 3. Toward the SLAU type schemes
4. The implicit time integration scheme for all Ma
5. Future challenges

2

Background

- As all real fluids are compressible, the compressible Navier-Stokes equation is theoretically a governing equation for flows at all Ma(Mach number) and sound propagation, however, there were problems to apply to the low Ma flows.
- The Godunov-type CFD schemes for compressible flows, in which the Riemann flux function computes the flux at the cell interface, are fundamental for modern high-resolution CFD methods used as the workhouse for the aerospace engineering.
- Recent advancements in the all-speed Riemann flux functions make it possible to compute flows in a broad range of Ma and to cover most aerospace applications.
- SLAU(Simple Low Dissipative Advection Upstream Splitting Method) type schemes we have developed are widely used as the all-speed Riemann flux scheme due to their robustness, simplicity and versatility.

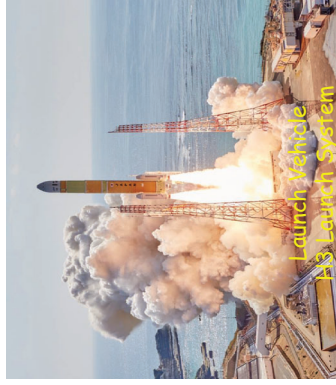
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Overview

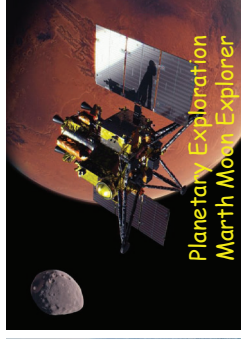
- The theories of the SLAU type schemes are presented along the history of progress from my point of view.
- I will emphasize the physical images of the schemes. It was helpful to improve the schemes.
- Application of the schemes at JAXA are shown for illustrating the use of SLAU type schemes and its wide range of applications. This also serves as an overview of JAXA's fluid-dynamic and aerodynamic researches.
(There are many other applications those don't use these schemes, of course.)
- Note that the simplified history I will talk about is not a whole history. There have been many excellent researches in this field, e.g. preconditioning methods, that I can't mention.

4

About JAXA



Launch Vehicle
H3 Launch System



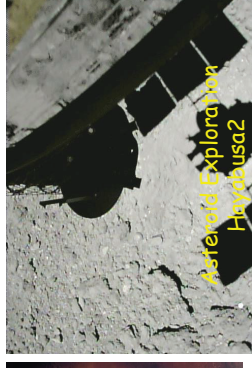
Planetary Exploration
Mars Moon Explorer



Human Space Program
HTV(H2 Transfer Vehicle)



Lunar Lander
(Next Generation SSD)



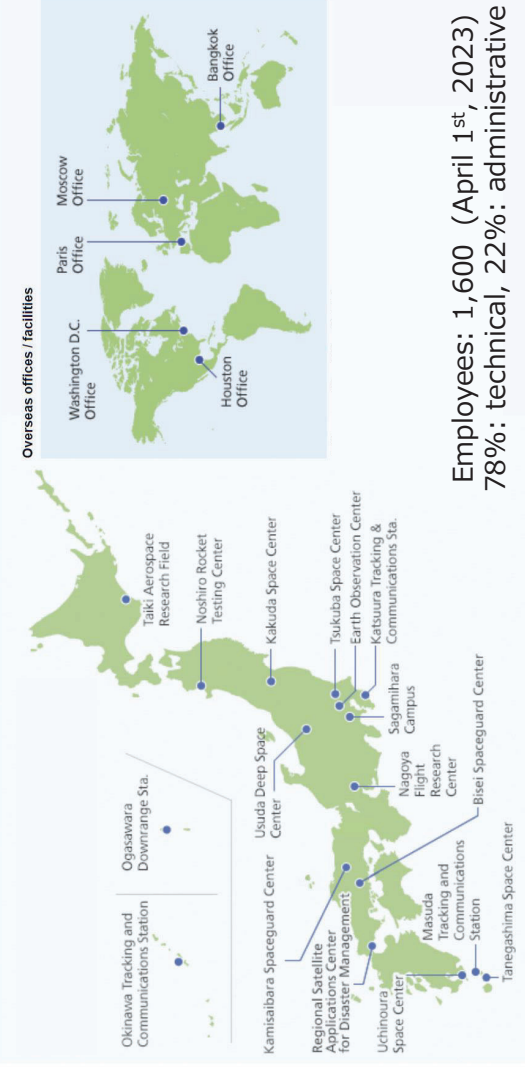
Asteroid Exploration
Hayabusa2



Launch Vehicle
Epsilon Launch System

5

Field Centers

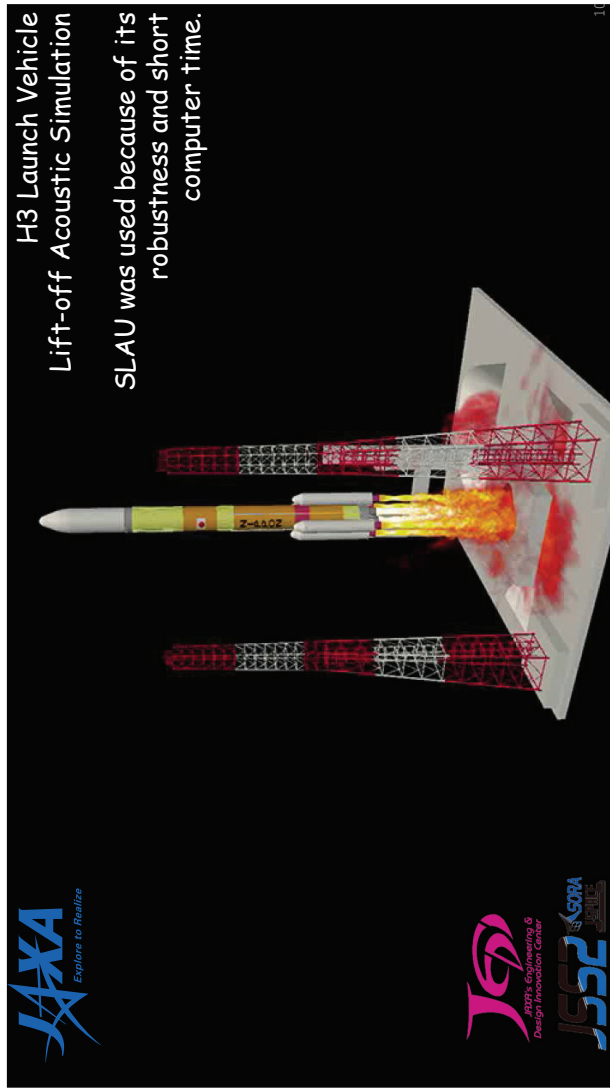


Employees: 1,600 (April 1st, 2023)
78%: technical, 22%: administrative



Applications of SLAU type schemes in JAXA's research projects

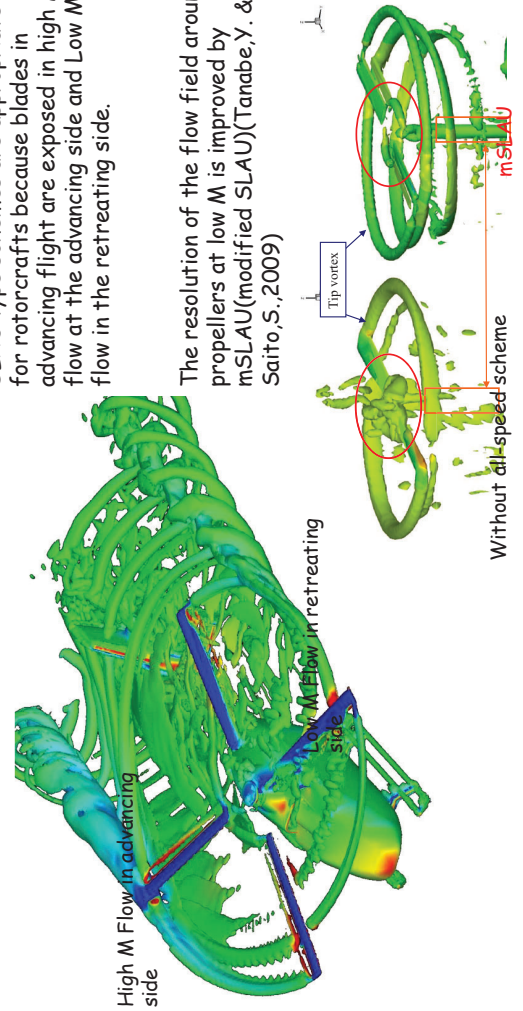
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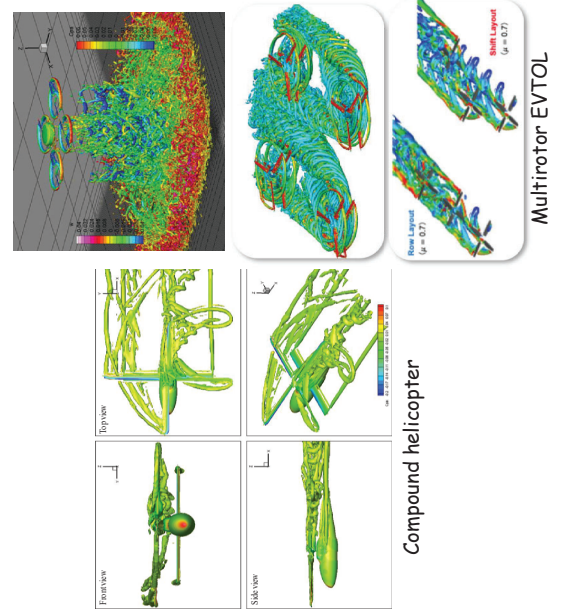
Application of mSLAU for Rotorcrafts(1)

SLAU type schemes are appropriate for rotorcrafts because blades in advancing flight are exposed in high Ma flow at the advancing side and Low Ma flow in the retreating side.

The resolution of the flow field around propellers at low M is improved by mSLAU(modified SLAU)(Tanabe,Y. & Saito,S.,2009)

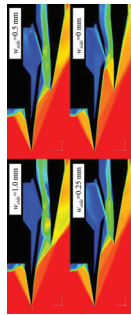


Application of mSLAU for Rotorcrafts(2)



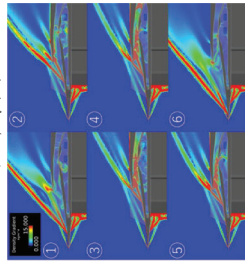
Applications of SLAU type schemes(3)

Hypersonic Intake



Scram Jet Engine

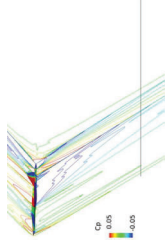
Numerical Simulation of the Side-Clearance Effect on the Supersonic Air-Inlet Performance for High Mach Integrated Control Experiment "HEMICO"
M. Ogasawara, S. Sano, H. Idekazu, Y. Yoshida, S. H. Wakabayashi, T. Kohno, T. Tetsuya Sato
Faculty of Engineering, and Tokonawa Campus (2018)



Pre-Cooled Turbo Tet Engine

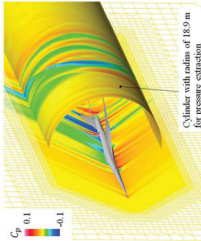
3-Dimensional Numerical Simulation of
Hyper sonic Intake for Pre-Coded Turbo Jet Engine
By Masatazu SANO, Toshiya YOKOI, Hidekazu YOSHIDA,
Tetsuya SATO, and Hideyuki TAGUCHI(2019)

Next Generation SST



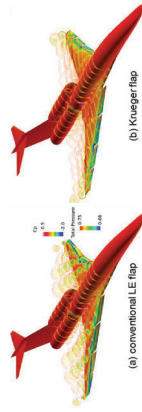
Sonic Boom Wave form

Numerical Simulation of Diffracted U-Shaped Sonic Boom Waveform
Masashi Karamori, Hiroaki Ishikawa, Yusuke Naka, and Yoshikazu Makino (20 22)



Low Boom Design

Low Boom Design
Robust Low-Boom Design in Primary Beam Carpet
Atsushi Ueno and Yoshihisa Makino

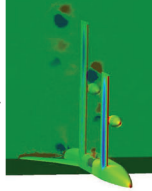


High Lift Devices

PRELIMINARY STUDY OF THE AERODYNAMIC CHARACTERISTICS OF KRUEGER FLAP ON THE SST CONFIGURATION
Dongyuan Kwak*, Keunho Ohno(2018)

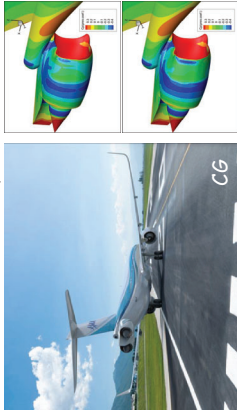
Applications of SLAU type schemes(4)

Electric/Hybrid Aircraft



UAV eVTOL

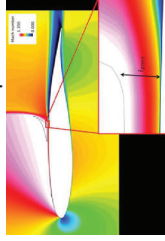
Forward Wing Tip Design for Induced Drag Reduction of Quod
Tilt Wing VTOL Un-Manned Plane in Cruise
Yoshitaka Aoki and Koji Murakami(2019)



Optimization of Electric Fan Caul

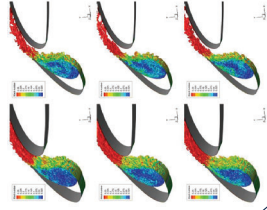
of Hybrid Aircraft

Transonic Transport Aircraft



Transonic Buffet

Transient Buffet Simulation over Supercritical Airfoil by
Unsteady FastAR Code
Takashi Ishida¹, Kazuhiro Tanihara², Atsushi Hashimoto, Takashi
Aoyama (2016)



(c) *annealed gold*

by Mark machines $a \in \text{Op}$, $a \in 1$ (middle), and $a \in 10$ (bottom).

Aeroacoustics of low Ma flow

Sat Noise Simulation on Unstructured Grid with Reduced Dissipation Approach
Ryota Sasaki, Takashi Ishida, Mitsuhiro Murayama, Yasuhide Ito, and Kazumi Yamamoto (2019)

FVM(Finite Volume Method) for Navier-Stokes Equation

NS Equation with the Source term in the Static Flame

$\iiint_{\Omega} \underline{Q}_i dV$	Conservative Variable	$\iint_{\partial\Omega} \underline{E} ds$	advection	$-\iint_{\partial\Omega} \underline{D} ds$	Diffusion	$-\iiint_{\Omega} \underline{S}_i dV$	Source
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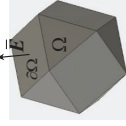
Euler Equation

$$\iiint_{\mathcal{V}} \underline{Q}_t dV + \iint_{\mathcal{S}} \underline{E} ds = 0$$

FVM Notation

$$\frac{1}{\Delta t} \Delta \bar{Q}_i + \frac{1}{V_i} \sum (\bar{E}_{i,j} \Delta s) = 0$$

$$MP(\bar{\mathcal{O}}_{-l+n}, \bar{\mathcal{O}}) \iiint \equiv \bar{\mathcal{O}}^\nabla$$



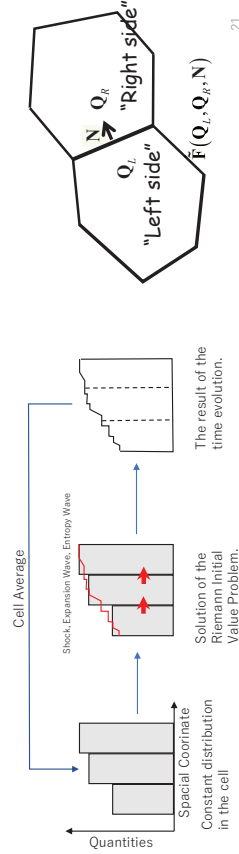
FVM is exact if the face average is exact.

$$\bar{\mathbf{E}}_{i,j} = \frac{1}{\Delta t \Delta S_i} \int_{n\Delta t}^{(n+1)\Delta t} \left(\iint_{\partial\Omega_{i,j}} \mathbf{E} ds \right) dt$$

However, obtaining the exact average flux is as difficult as solving Euler equation exactly. So, we need approximations.

What is the Riemann Flux Function?

- Godunov Scheme (Godunov, S., 1959)
 - Assuming the constant distribution in the cell (control volume), flux can be computed as the solution to the Riemann Initial Value problems.
- The Riemann Flux Function
 - The detail of the solution will be lost only the influence of the flux at the interface will remain. The flux is called the Riemann flux.
 - I will call the functions that give "appropriate" flux from two independent physical states "The Riemann Flux Function" or "The Riemann Flux".



21

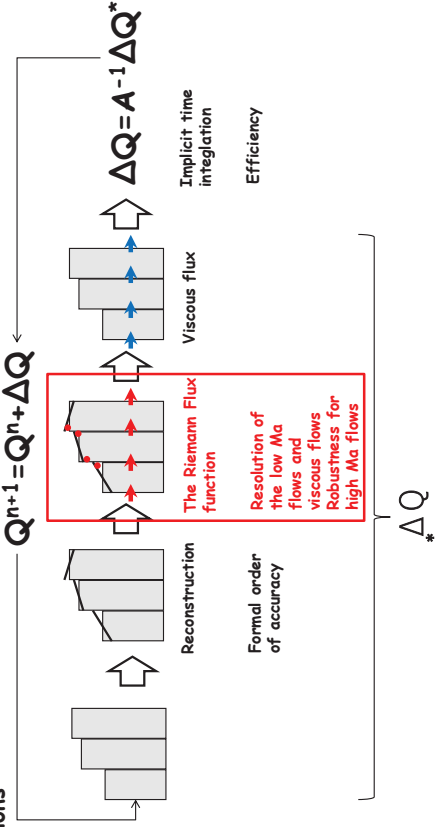
Higher Order Godunov-type Method

- $O(\Delta x)$ discontinuity at the interface produce the 1st order error.
- Higher order accuracy by the reconstruction in the cell can be achieved.
 - The distribution of the physical value is reconstructed by some method before the Riemann flux computation. The error at the interface will be reduced.
 - MUSCL (Monotonic Upstream-centered Scheme for Conservation Laws) (van Leer, B., 1979)
- Many methods have been developed and studied, and some techniques have become standard, such as MUSCL, TVD, ENO, WENO, etc. The details are omitted in this talk.

22

Higher Efficiency is Achieved by the Implicit Algorithm

The Procedure of the Implicit CFD Method for the Compressible Navier-Stokes Equations



23

Remarks on the Riemann Fluxes

- The exact solution has little benefit.
 - Although the exact solution can be computed, it is merely an approximation in multidimensional flows or higher-order schemes.
 - The theory is simple, however, the computation is time consuming.
 - The numerical dissipation is too big for low Ma flows.
 - This tends to create anomalies at the shock wave known as the carbuncle phenomena.
- Therefore approximate Riemann fluxes are more attractive.

24

History and Theories

2. Toward the AUSM type Schemes

Common Form of AUSM type schemes

Flux of the Euler equation in 3-D FVM form can be written as:

$$\tilde{\mathbf{E}} = \dot{m} \mathbf{\Phi} + p \mathbf{N} \quad \mathbf{\Phi} = \begin{pmatrix} 1 \\ u \\ v \\ w \\ h \end{pmatrix} \quad \mathbf{N} = \begin{pmatrix} 0 \\ x_n \\ y_n \\ z_n \\ 0 \end{pmatrix}$$

Concentration Face normal

$$\dot{m} = \rho V_n \quad h = \frac{e+p}{\rho}$$

Mass flux Total Enthalpy

The flux of AUSM(Liou, M.-S., Steffen Jr, C.J., 1993) type schemes is close to the definition of the Euler flux.

$$\tilde{\mathbf{E}} = \frac{\dot{m} + |\dot{m}|}{2} \mathbf{\Phi}_L + \frac{\dot{m} - |\dot{m}|}{2} \mathbf{\Phi}_R + \tilde{p} \mathbf{N}$$

The mass Flux $\dot{m}$ and the interface pressure $\tilde{p}$ are everything for the AUSM type scheme. I will show how these schemes were made.

CIR(Courant,R., Isaacson,E. & Rees,M. 1952) scheme

The Godunov scheme is the 1st order upwind difference for the Euler Equation.

Here we think about the spatially one-dimensional linear hyperbolic system.

$$\mathbf{Q}_t + \mathbf{A} \mathbf{Q}_x = 0$$

CIR scheme which is 1st order upwind scheme can be written in the FVM form.

$$\mathbf{Q}^{n+1} = \mathbf{Q}^n - \frac{\Delta t}{\Delta x} (\tilde{\mathbf{E}}_{i+1/2} - \tilde{\mathbf{E}}_{i-1/2})$$

$$\tilde{\mathbf{E}}_{i+1/2} = \frac{1}{2} \{ (\mathbf{A} \mathbf{Q})_i + (\mathbf{A} \mathbf{Q})_{i+1} - |\mathbf{A}| (\mathbf{Q}_{i+1} - \mathbf{Q}_i) \}$$

Here we use the following orthogonal transformation.

$$\mathbf{A} = \mathbf{T} \text{diag}(\lambda_k) \mathbf{T}^{-1}$$

$$|\mathbf{A}| = \mathbf{T} \text{diag}(|\lambda_k|) \mathbf{T}^{-1}$$

Two origins of the approximate Riemann Fluxes: FDS and FVS

CIR

$$\tilde{\mathbf{E}} = \frac{1}{2} \{ (\mathbf{A} \mathbf{Q})_L + (\mathbf{A} \mathbf{Q})_R - |\mathbf{A}| (\mathbf{Q}_R - \mathbf{Q}_L) \}$$

$$\mathbf{E} = \mathbf{A} \mathbf{Q} \quad \text{Equivalent in linear case} \quad \mathbf{A}^\pm = \frac{\mathbf{A} \pm |\mathbf{A}|}{2}$$

$$\tilde{\mathbf{E}} = \mathbf{A}^+ \mathbf{Q}_L + \mathbf{A}^- \mathbf{Q}_R$$

Extension for non-linear equation

$$\tilde{\mathbf{E}} = \frac{1}{2} \{ \mathbf{E}_L + \mathbf{E}_R - |\tilde{\mathbf{A}}| (\mathbf{Q}_R - \mathbf{Q}_L) \}$$

$$\tilde{\mathbf{E}} = \frac{1}{2} \{ \mathbf{E}_L + \mathbf{E}_R - |\tilde{\mathbf{A}}| (\mathbf{Q}_R - \mathbf{Q}_L) \}$$

$\tilde{\mathbf{A}}$: from Roe Averaged Value

FVS(Flux Difference Splitting) (Roe,P.L. 1981)

Drawbacks: Strong shock and expansion

FVS(Flux Vector Splitting)

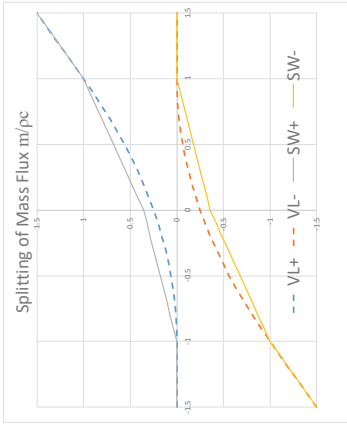
(Steger,J.L.&Warming,R.F, 1979)

Drawbacks: Large error at contact surface

Improvement of FVS(van Leer,B.,1982)

Challenge: Avoiding glitches in SW-FVS

- Smooth transition at switching using polynomials is introduced.
- Simpler form than SW-FVS was achieved.

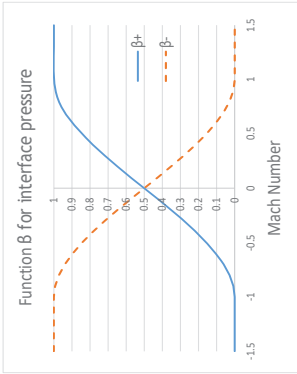


Improvement of FVS (cont'd)

By separating the flux to the advection and the pressure term, it was found that interface pressure can be written in the weighted average of the pressure on the left and right.

$$\tilde{p} = p_L^+ + p_R^- = p_L \beta_L^+(M_L) + p_R \beta_R^-(M_R)$$

Although it became clear by Hänel's modification, it has been included in the van Leer's FVS. We will keep using this interface pressure in AUSM type schemes.



Improvement of FVS (Hänel & Schwane,1989)

Challenge : Preservation of total enthalpy which is crucial in high Ma flows

The vL-FVS can be rewritten to separate contributions of advection and pressure.

$$_{vL} \tilde{\mathbf{E}} = \tilde{\mathbf{E}}_{vL}^+ + \tilde{\mathbf{E}}_{vL}^- \quad \tilde{\mathbf{E}}_{vL}^{\pm} = \dot{m}^{\pm} \Phi^{\pm} + p^{\pm} \mathbf{N}$$

$$\Phi^{\pm} = \Phi_L + \left(0, 0, 0, -\frac{(V_{rel} - c_L)^2}{\gamma + 1} \right)^T \quad \Phi^{\pm} = \Phi_R + \left(0, 0, 0, -\frac{(V_{rel} + c_R)^2}{\gamma + 1} \right)^T$$

Preservation of total enthalpy is improved by small change in the energy term.
→Hänel's FVS

$$_{Hanel} \tilde{\mathbf{E}} = \tilde{\mathbf{E}}_L^+ + \tilde{\mathbf{E}}_R^-$$

$$\tilde{\mathbf{E}}_{Hanel}^{\pm} = \dot{m}^{\pm} \Phi^{\pm} + p^{\pm} \mathbf{N}$$

Very close form to the definition of the Euler flux.

From FVS to AUSM

Challenge: Accuracy for contact surfaces and viscous flows

AUSM is well known, however, here we explain using the concept of SFS(Simplified Flux Vector Splitting).(Jounouchi,T. et al., 1991)
Hänel's FVS can be written as follows.

$$_{Hanel} \tilde{\mathbf{E}} = \quad_{vL} \dot{m}_L^+ \Phi_L^+ + \quad_{vL} \dot{m}_R^- \Phi_R^- + \tilde{p} \mathbf{N}$$

In this form, particles from each side transport the concentration of each side independently. This works as numerical molecular diffusion and is not the correct behavior of the Euler flux. If the fluid as a particle transports the concentration, one particle on the upwind side should do.

$$_{SFS} \dot{m} = \dot{m}_L^+ + \dot{m}_L^- \quad \text{with a green oval containing } (\dot{m}_L^+ + \dot{m}_L^-) \Phi^+$$

$$_{AUSM-F} \tilde{\mathbf{E}} = \frac{\dot{m} + |\dot{m}|}{2} \Phi_L + \frac{\dot{m} - |\dot{m}|}{2} \Phi_R + \tilde{p} \mathbf{N}$$

The common form of the AUSM type schemes is obtained.

Common Form of AUSM type schemes

$$\tilde{\mathbf{E}} = \frac{\tilde{m} + |\tilde{m}|}{2} \Phi_L + \frac{\tilde{m} - |\tilde{m}|}{2} \Phi_R + \tilde{p} \mathbf{N}$$

The mass flux and the interface pressure are everything for the AUSM type scheme.

History and Theories

3. Toward the All-Speed Schemes (1)Interface Pressure

The interface pressure also works as the numerical dissipation and needs to be adjusted for low M flows.

Numerical Dissipation in the Interface Pressure

AUSM type schemes inherit the interface pressure from van Leer 's FVS

$$\tilde{p} = \beta^+ p_L + \beta^- p_R$$

Exactly rewritten as;

$$\tilde{p} = \frac{p_L + p_R}{2} + \frac{\beta^+ - \beta^-}{2} (p_L - p_R) + (\beta^+ + \beta^- - 1) \frac{p_L + p_R}{2}$$

$M \ll 1 \rightarrow \frac{3}{4\gamma} \frac{\rho_L + \rho_R}{2} (\tilde{c} V_{nL} - V_{nR})$

This term works as numerical bulk viscosity with the scale of the sound speed. The scale is good only for the flow at near-sound speed.

Controlling the Numerical Dissipation

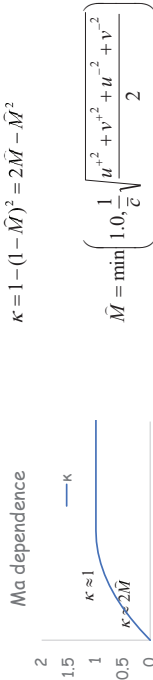
Control the scale of dissipation by smooth function of Ma.

$$\tilde{p} = \frac{p_L + p_R}{2} + \frac{\beta^+ - \beta^-}{2} (p_L - p_R) + (\beta^+ + \beta^- - 1) \frac{p_L + p_R}{2}$$

(Shima E., 2006)

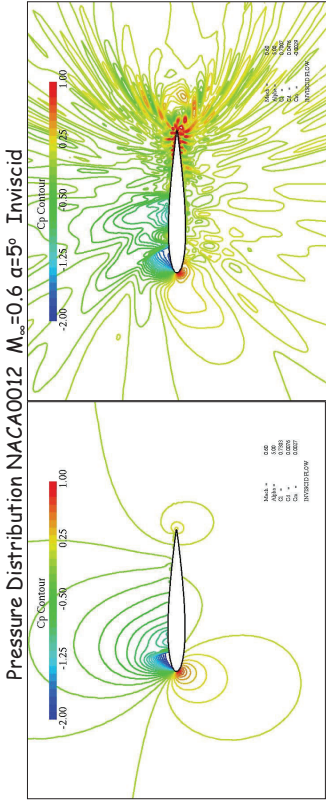
$$\tilde{p} = \frac{p_L + p_R}{2} + \frac{\beta^+ - \beta^-}{2} (p_L - p_R)$$

(Ma)x(Sound Speed) = Advective Speed



Choice of the interface Ma

The use of the multidimensional Mach number gives good solution.



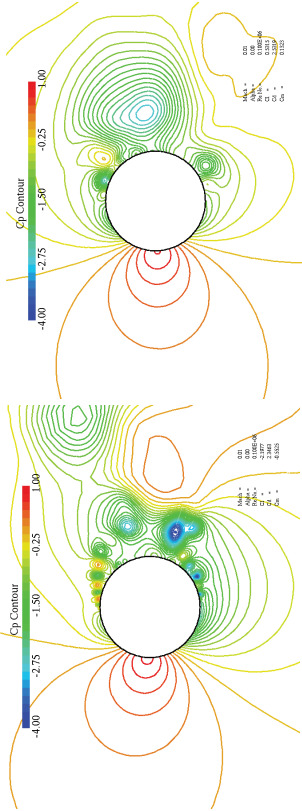
Multidimensional Mach Number

$$\tilde{M} = \min \left(1.0, \frac{1}{\tilde{c}} \sqrt{\frac{u^2 + v^2 + w^2}{2}} \right)$$

One Dimensional Mach Number

$$\tilde{M} = \min \left(1.0, \frac{1}{\tilde{c}} \sqrt{\frac{U^2 + V^2}{2}} \right)$$

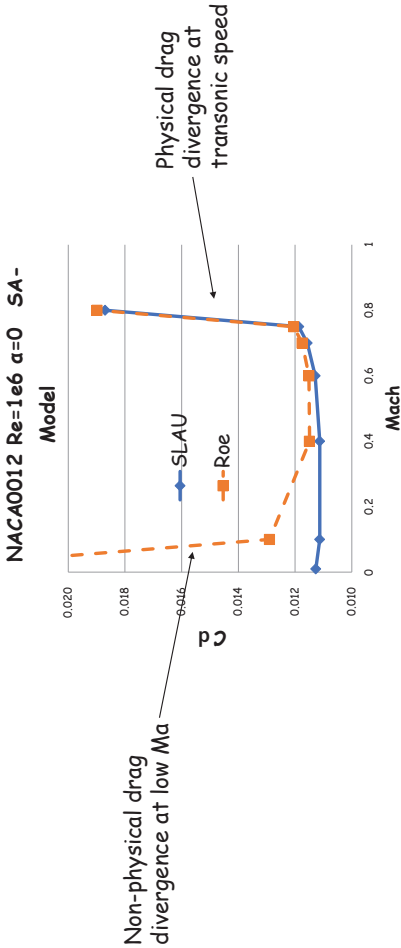
Improved Resolution for Viscous Flows M_∞=0.01 Re=10⁶



SLAU 3rd order without limiter
MFGS implicit scheme

Roe scheme 3rd order without limiter
MFGS implicit scheme

Non-physical drag divergence at low Ma was corrected by SLAU.



SLAU2 for hypersonic and general EOS

AUSM

$$\begin{aligned} \tilde{p} &= \beta^+ p_L + \beta^- p_R \\ &= \frac{p_L + p_R}{2} + \frac{\beta^+ - \beta^-}{2} (p_L - p_R) + (\beta^+ + \beta^- - 1) \frac{p_L + p_R}{2} \end{aligned}$$

Change the scale of the
dissipation at Low Ma.

$$\tilde{p} = \frac{p_L + p_R}{2} + \frac{\beta^+ - \beta^-}{2} (p_L - p_R) + \kappa \frac{\beta^+ + \beta^- - 1}{2} \frac{p_L + p_R}{2}$$

Too small for the hypersonic
speed

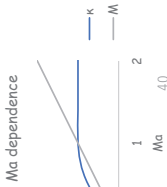
It only work as proper
dissipation when $p \propto \rho c^2$

SLAU
(Shima,E,2008)

SLAU2

(Kiamura,K. &
Shima,e,2011)

$$\tilde{p} = \frac{p_L + p_R}{2} + \frac{\beta^+ - \beta^-}{2} (p_L - p_R) + \tilde{M} \frac{\beta^+ + \beta^- - 1}{2} \tilde{\rho} \tilde{c}^2$$



History and Theories

3. Toward the All-Speed Schemes (2) Mass Flux

The pressure difference term in the mass flux works as the dissipation for the acoustic wave.
But it also causes the carbuncle phenomena.

41

Utilize the Mass Flux of FDS

They were natural choices. However, the mass flux of FDS is better to control term by term.

SHUS(Simple High-resolution Upwind Scheme) (Shima,E. et al., 1994)

$$\dot{m} = \frac{1}{2} \left\{ (\rho V_n)_L + (\rho V_n)_R - |\bar{V}_n| \Delta \rho - \frac{|\bar{M} + 1| - |\bar{M} - 1|}{2} \frac{|\bar{M} + 1| + |\bar{M} - 1| - 2|\bar{M}|}{2\bar{c}} \Delta p \right\}$$

$$\Delta q \equiv q_R - q_L$$

For very low $\bar{M}$, the mass flux approaches to:

$$\bar{M} \rightarrow 0$$

$$\dot{m} = \frac{1}{2} \left\{ (\rho V_n)_L + (\rho V_n)_R - |\bar{V}_n| \Delta \rho - \frac{1}{\bar{c}} \Delta p \right\}$$

For acoustic wave, the last term works as dissipation.

$$\Delta p = \bar{c}^2 \Delta \rho$$

$$\dot{m} = \frac{1}{2} \left\{ (\rho V_n)_L + (\rho V_n)_R - (\bar{c} + |\bar{V}_n|) \Delta \rho \right\}$$

43

Early AUSM type schemes inherit the mass flux of FVS.

FVS

$$FVS \dot{m} = g^+(M_L) \rho_L c_L + g^-(M_R) \rho_R c_R$$

SFS

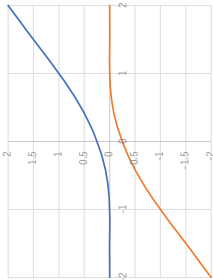
$$SFS \dot{m} = g^+(M_L) \rho_L \bar{c} + g^-(M_R) \rho_R \bar{c}$$

AUSM

$$\bar{M} = g^+(M_L) + g^-(M_R)$$

$$AUSM \dot{m} = \frac{\bar{M} + |\bar{M}|}{2} \rho_L \bar{c} + \frac{\bar{M} - |\bar{M}|}{2} \rho_R \bar{c}$$

$$g^\pm(M) = \begin{cases} \pm \frac{1}{4} (M \pm 1)^2 & \text{if } |M| < 1, \\ \frac{1}{2} (M \pm |M|) & \text{otherwise} \end{cases}$$



AUSM-DV

(Wada,Y.,1997)

$$AUSM-DV \dot{m} = \rho_L \left\{ \alpha_L \frac{V_{nd} + |V_{nd}|}{2} + (1 - \alpha_L) g^+(M_L) c_L \right\} + \rho_R \left\{ \alpha_R \frac{V_{nr} - |V_{nr}|}{2} + (1 - \alpha_R) g^-(M_R) c_R \right\}$$

42

Effect of the Pressure Difference Term

Here we will see the effect of the pressure difference term.

SHUS

$$\dot{m} = \frac{1}{2} \left\{ (\rho V_n)_L + (\rho V_n)_R - |\bar{V}_n| \Delta \rho - \frac{|\bar{M} + 1| - |\bar{M} - 1|}{2} \frac{|\bar{M} + 1| + |\bar{M} - 1| - 2|\bar{M}|}{2\bar{c}} \Delta p \right\}$$

The Δp term in the mass flux will cause the carbuncle at shockwave.

(Liou,M.-S.,2000)

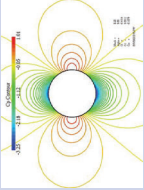
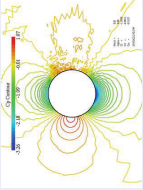
ST(Simplest)-AUSM(Shima,E.,2000)

$$\dot{m} = \frac{1}{2} \left\{ (\rho V_n)_L + (\rho V_n)_R - |\bar{V}_n| \Delta \rho \right\}$$

No Pressure Difference Term

44

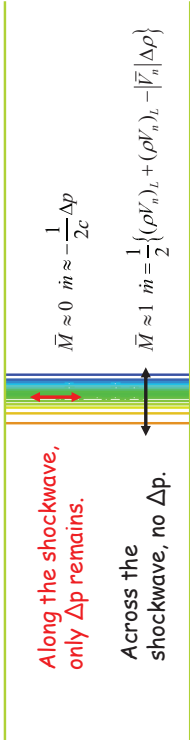
Both schemes work well in general.
The Δp term can eliminate noise in low Ma but cause the carbuncle in supersonic.

	Δp	$M > 1$	$M < 1$
SHUS	ON	 Carbuncle	 Good
ST-AUSM	OFF	 Good No carbuncle	 NG

Review of Mass Flux of SHUS at the Shock Wave

Review of the Δp term of SHUS at the shock wave.

$$\dot{m} = \frac{1}{2} \left\{ (\rho V_n)_L + (\rho V_n)_R - |\bar{V}_n| \Delta p - \frac{|\bar{M} + 1| - |\bar{M} - 1|}{2} \bar{\rho} \Delta V_n - \frac{|\bar{M} + 1| + |\bar{M} - 1| - 2|\bar{M}|}{2\bar{c}} \Delta p \right\}$$



The Δp term along the shockwave causes the carbuncle phenomena.

Mass Flux for SLAU

We will see the effect of the pressure difference term.

SHUS

$$\dot{m} = \frac{1}{2} \left\{ (\rho V_n)_L + (\rho V_n)_R - |\bar{V}_n| \Delta p - \frac{|\bar{M} + 1| - |\bar{M} - 1|}{2} \bar{\rho} \Delta V_n - \frac{|\bar{M} + 1| + |\bar{M} - 1| - 2|\bar{M}|}{2\bar{c}} \Delta p \right\}$$

$$\Delta q \equiv q_R - q_L$$

ST-AUSM

$$\dot{m} = \frac{1}{2} \left\{ (\rho V_n)_L + (\rho V_n)_R - |\bar{V}_n| \Delta p \right\}$$

SLAU(Shima,E,2008)

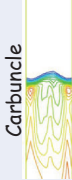
$$\dot{m} = \frac{1}{2} \left\{ (\rho V_n)_L + (\rho V_n)_R - |\bar{V}_n| \Delta p - \frac{1}{c} \Delta p \right\}$$

Simply eliminates in whole high Ma region.

$$1 - \kappa = (1 - \bar{M})^2 \quad \bar{M} = \min \left(1.0, \sqrt{\frac{1}{2} \frac{u'^2 + v'^2 + u'^2 + v'^2}{c^2}} \right)$$

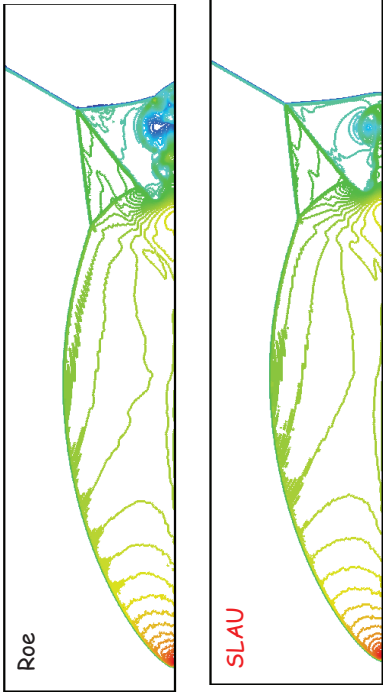
The final mass flux of SLAU includes the term $\frac{1}{2} [(\rho V_n)_L + (\rho V_n)_R - |\bar{V}_n| \Delta p - (1 - \kappa) \frac{1}{c} \Delta p]$ for strong expansion, but 1 unit from $g = \max(-\min(\rho V_n, 0), \min(\rho V_n, 0))$

SLAU is good in both low and high Ma.

	Δp		$M > 1$	$M < 1$
	High M	Low M		
SHUS	ON	ON	 Carbuncle	 Good
ST-AUSM	OFF	OFF	 Good No carbuncle	 NG
SLAU	OFF	ON	 Good No carbuncle	 Good

The carbuncle phenomena are more harmful in complex flow fields.

Double Mach Reflection



Improved mass flux of SD-SLAU(Shima,E. et al., 2010)

The mass flux of SD(Shock Detecting)-SLAU is defined as;

$$\dot{m} = \frac{1}{2} \left\{ \rho^+ V_n^+ + \rho^- V_n^- - |\bar{V}_n| \Delta p - \theta \frac{(\Delta p)}{p} \left[\frac{|\bar{M}| + 1 + |\bar{M}| - 1 - 2|\bar{M}|}{2\bar{c}} \right] \Delta p \right\}$$

Switching function θ is designed to be;

$$\theta \left(\frac{\Delta p}{p} \right) \begin{cases} \approx 0 & \text{in parallel direction to the shock} \\ \approx 1 & \text{otherwise} \end{cases}$$

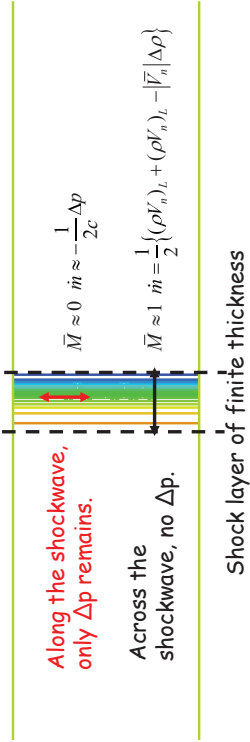
Basic mass flux 2 : SLAU

$$\dot{m} = \frac{1}{2} \left\{ \rho^+ V_n^+ + \rho^- V_n^- - |\bar{V}_n| \Delta p - \frac{\bar{c}}{\bar{c}} \Delta p \right\}$$

Basic mass flux 1 : SHUS

$$\dot{m} = \frac{1}{2} \left\{ \rho^+ V_n^+ + \rho^- V_n^- - |\bar{V}_n| \Delta p - \frac{|\bar{M}| + 1 + |\bar{M}| - 1}{2} \bar{\rho} \Delta V_n - \frac{|\bar{M}| + 1 + |\bar{M}| - 1 - 2|\bar{M}|}{2\bar{c}} \Delta p \right\}$$

The Shock Layer is a Fictitious Space



As the behavior of these pressure differences are those of the Euler equation, they must be physically correct.

However, the shock layer of finite thickness is not a physical space. Computing correctly in the fictitious space seems to cause problem.

Therefore better scheme will be achieved, if the pressure difference is eliminated only in the shock layer.

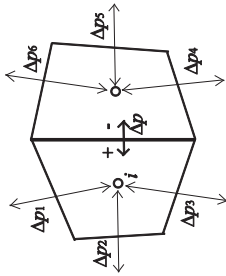
Shock Detecting Function

Actual form of fp is;

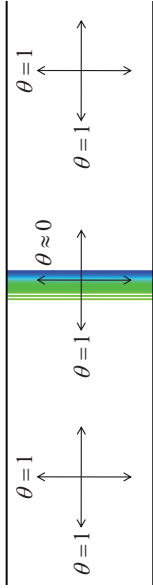
$$f_p(\Delta p / p) = \min \left(1, \left(\frac{C_{SD2} |\Delta p| / \bar{p} + C_{SD1}}{|\Delta p|_{\max} / \bar{p} + C_{SD1}} \right)^2 \right)$$

$$C_{SD2} \approx 10$$

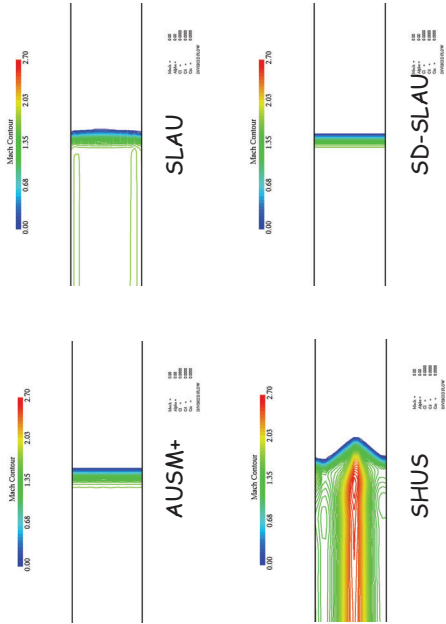
$$C_{SD1} \approx 0.1$$



Shock



1D shock propagation in distorted 2D channel with distorted mesh



Other Choices for Δp term

The general form of the SLAU type schemes

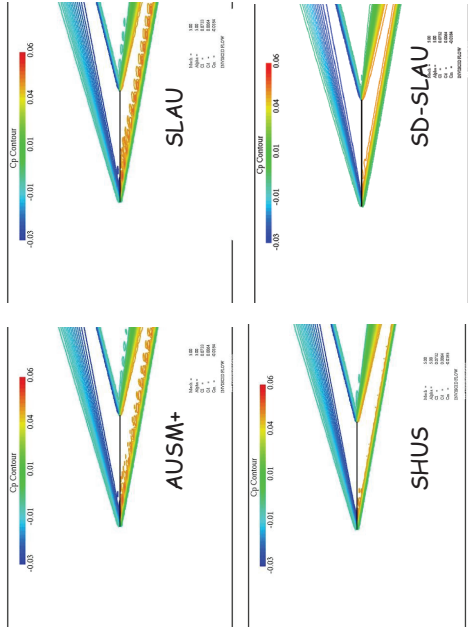
$$m_i = \frac{1}{2} \left\{ (\rho V)_n)_L + (\rho V)_n)_R - |\vec{V}|_n |\Delta \rho - \boxed{f_i} \frac{L}{c} \Delta p \right\}$$

scheme	f_p	Advantages	Drawback
SLAU	$1 - \kappa$		
SD-SLAU	$\min \left(1, \left(\frac{C_{\text{vis}} \Delta p / \bar{p} + C_{\text{vis}}}{ \Delta p _{\text{min}} / \bar{p} + C_{\text{vis}}} \right) \max(0.1, \vec{V}) \right)$	Eliminating noise in supersonic flows	More complex
No-name	$\frac{1 - \kappa}{\kappa + M_c^2}$	Monotonic even in low Ma (Ma > Mc)	CFL restriction Cut-off Ma Dissipation of sound
WS(Wiggle Sensing)-SLAU(Shima,E. et al., 2013)	$\left\{ 1 + \min \left(\frac{\max(\vec{\Delta x}_p , \vec{\Delta y}_p)}{\max(\vec{\Delta x}_p , \vec{\Delta y}_p , x)} \right) \frac{1 - \kappa}{\kappa + M_c^2} \right\} (1 - \kappa)$	Eliminating wiggles in very low Ma flows	CFL restriction Cut-off Ma More complex
UD(Uniformly Damping)-SLAU(Shima,E. et al., 2015)	$(1 - \kappa) \max \left(1, \frac{\overline{cT_c}}{\Delta x} \right)$	Eliminating unnecessary acoustic wave in low Ma	Needs GC-SMAC

We have better schemes for extreme conditions, however, don't have a simple and universal scheme yet.

SLAU is the simplest and works well in most aerodynamic flows.

Oblique shock around a thin plate $M=6$ $\alpha=5^\circ$



Time Integration Method for All Ma

Note on Time Integration Methods

- Time step restriction from CFL condition is severe for low Ma flows due to large ratio of the sound speed to the advection speed.
- The regular implicit time integration methods such as LU-SGS is effective for $Ma > 0.1$. Therefore such schemes are good enough for most aerodynamic flows
- The preconditioning methods as the implicit time integration work well in $Ma < 0.1$ for steady flows.
- However, the choice of cut-off Ma is difficult and they causes storing dissipation on the sound wave propagation.
- GC(Generalized Compressible)-SMAC(Simplified Marker and Cell Method)(Shima,E., etc.,2018), which smoothly approaches to SMAC for the incompressible flows in very low Ma, seems promising.

Convert R.H.S. residual to the entropy variable

Use the MUSCL FVM for conservative variables with all speed Riemann flux for R.H.S. evaluation.

$$\bar{Q}_i + \frac{1}{V_i} \sum_j (\tilde{\mathbf{E}}_{i,j} - \tilde{\mathbf{R}}_{i,j}) ds_{i,j} = 0$$

Use the working variables for L.H.S. evaluation for simplicity. Variation of working variables are is obtained by mapping;

$$\delta \mathbf{q}^* \equiv \frac{\partial \mathbf{q}}{\partial \mathbf{Q}} \left[-\delta t \frac{1}{V_i} \sum_j (\tilde{\mathbf{E}}_{i,j} - \tilde{\mathbf{R}}_{i,j}) ds_{i,j} \right]$$

After approximate matrix inversion in working variables, variation of conservative variables is obtained by reverse mapping;

$$\delta \mathbf{Q} \equiv \frac{\partial \mathbf{Q}}{\partial \mathbf{q}} \mathbf{B}^{-1} \delta \mathbf{q}^*$$

Focuses on the evaluation of the inviscid terms, since we treat the viscosity term in regular way,

The Equation become simple in the entropy variables.

Use entropy variables (velocity, pressure, entropy) as working variables. Compressible NS equations using them are written as;

$$\mathbf{u}_t + \mathbf{u} \cdot \nabla \mathbf{u} + \frac{1}{\rho} \nabla p + \mathbf{R}_u = 0$$

$$p_t + \mathbf{u} \cdot \nabla p + \rho c^2 \nabla \cdot \mathbf{u} + R_p = 0$$

$$s_t + \mathbf{u} \cdot \nabla s + R_s = 0$$

Semi-discretize in time to the δ form using the first order implicit method.

$$\delta \mathbf{u} + \delta t \left\{ \mathbf{u} \cdot \nabla \delta \mathbf{u} + \frac{1}{\rho} \nabla \delta p + \frac{\partial \mathbf{R}_u}{\partial \mathbf{u}} \delta \mathbf{u} \right\} = \delta \mathbf{u}^*$$

$$\delta p + \delta t \left\{ \mathbf{u} \cdot \nabla \delta p + \rho c^2 \nabla \cdot \delta \mathbf{u} + \frac{\partial R_p}{\partial p} \delta p \right\} = \delta p^*$$

$$\delta s + \delta t \left\{ \mathbf{u} \cdot \nabla \delta s + \frac{\partial R_s}{\partial s} \delta p \right\} = \delta s^*$$

Operator splitting of least error create SMAC like scheme

L.H.S operator is splitting in the least error manner. The linear equation for the pressure can be solved by similar way as SMAC method for incompressible CFD.

$$\frac{\partial \mathbf{Q}}{\partial \mathbf{q}} \left[\mathbf{I} + \delta t' \left\{ \mathbf{A}_u + \mathbf{A}_p + \mathbf{A}_s \right\} \right] \frac{\partial \mathbf{q}}{\partial \mathbf{Q}} \Delta \mathbf{Q} = \mathbf{Q} - \theta \mathbf{Q}^* + \theta_2 \mathbf{Q}^{*1} - \delta t' \frac{1}{\Omega_i} \sum_j (\tilde{\mathbf{E}}_{i,j} - \tilde{\mathbf{R}}_{i,j}) dS_{i,j}$$

Advection Pressure

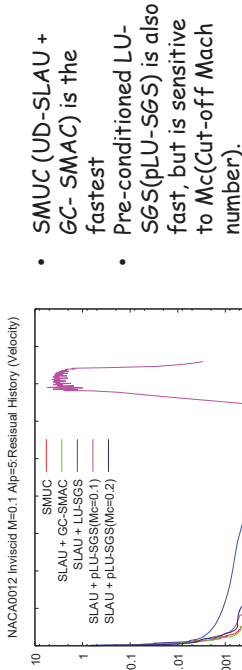
$$(\theta_1, \theta_2) = \left(\frac{2\theta+2}{\theta+2}, \frac{\theta}{\theta+2} \right)$$

$$\delta t' = \frac{2}{\theta+2} \delta t$$

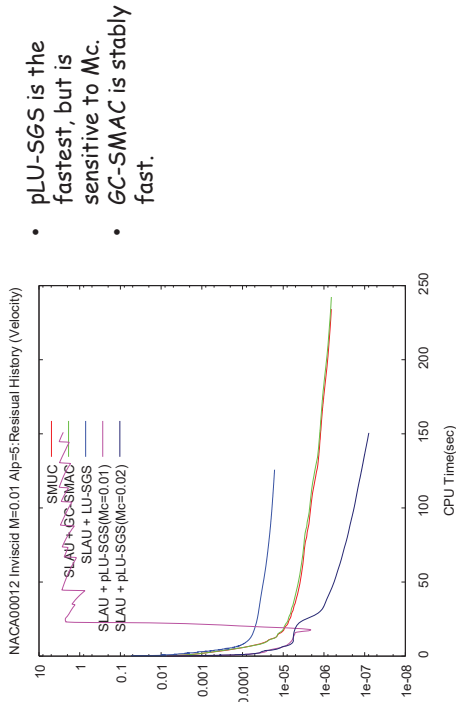
$$\theta = \begin{cases} 0 & : \text{First order accuracy} \\ 1 & : \text{Second order accuracy} \end{cases}$$

Applying this operator splitting, the pressure part can be solved by the similar method as SMAC.

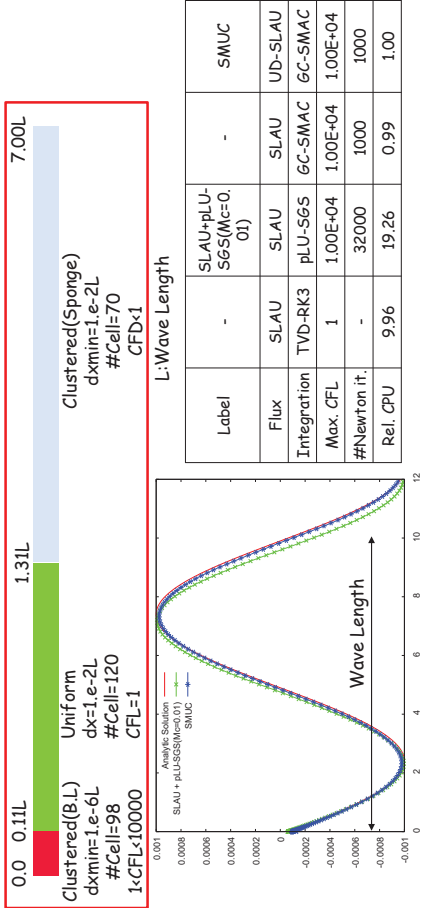
Inviscid flow at $M = 0.1$ around NACA0012
Convergence history of velocity residuals



Inviscid flow at $M = 0.01$ around NACA0012
Convergence history of velocity residuals



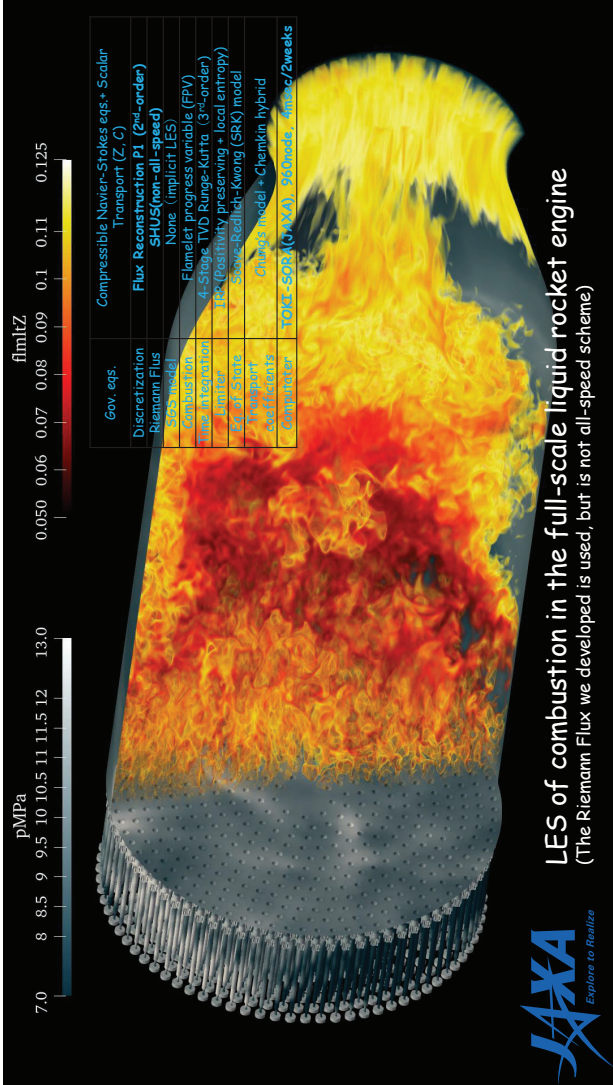
1D Sound Propagation Simulating Sound Emitted From B.L. ($CFL_{max}=10^4$)



SMUC is 10 times faster than the explicit scheme and 19 times faster than the time derivative pre-conditioned scheme (pLU-SGS) with $Mc=0.01$.
The result of pLU-SGS show small deviation even after 32000 Newton iterations.

Future Challenges

- Robustness
 - SLAU works fine with FR for LES alone, however, breakdowns with the combustion. On the other hand, non all-speed SHUS works. We don't know why and how to fix.
- Accuracy
 - Utilizing the flexibility of the SLAU type scheme, such as choice of the some averaged values in the scheme or choice of the reconstruction scheme, new features like conservation of the kinetic energy or entropy might be possible.
- Efficiency
 - GC(Generalized Compressible)-SMAC works fine for very low Ma without adjustment of the cut-off Ma in the trial cases. As the present version is still primitive stage, future researches are expected.



Characteristics of the Riemann Fluxes

	Godunov (Exact Riemann Flux)	Roe (Approximate Riemann Flux)	FVS (Approximate Riemann Flux)	SLAU
Theory	Simple	Simple	Simple	Complex
Cost	NG -Iteration	NG -Orthogonal decomposition	Good	Good
Supersonic(Ma>2)	NG -Carbuncle	NG -Carbuncle -Breakup at strong expansion	Good	Good
Low Ma(<0.1)	NG -Too dissipative	NG -Too dissipative	NG -Too dissipative	Good
Moderate Ma (0.1<Ma<2)	Good	Good	Good	Good
Viscous flow	Good	Good	NG -Too dissipative	Good

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Remarks on compressible CFD schemes

- The problem of compressible CFD schemes on computation of low Ma flows have been almost solved.
- Large numerical dissipation.
- Solved by all-speed Riemann fluxes.
- Strict time step size restriction
 - In Ma > 0.1, popular and simple implicit time integration schemes, such as LU-SGS, work well.
 - For lower Ma, the preconditioning methods are effective.
 - Even in Ma<0.01, efficient methods have been constructed. (such as, GC-SMAC, Shima.E.,2015)
- Roundoff error due to small change.
- Can be solved by separating the derivations from the uniform values. In most aerospace applications, the double precision variables are good enough.

Thank you for your attention!